

Non-Abelian Zeta Function, Fokker-Planck Equation and Projectively Flat Connection

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Abstract

Over the moduli space of rank n semi-stable lattices is a universal family of tori. Along the fibers, there are natural differential operators and differential equations, particularly, the heat equations and the Fokker-Planck equations in statistical mechanics. In this paper, we explain why, by taking averages over the moduli spaces, all these are connected with the zeros of rank n non-abelian zeta functions of the field of rationals, which are known lie on the central line except a finitely many if $n \geq 2$. Certainly, when $n = 1$, our current work recovers that of Armitage, which from the beginning motivates ours. However, we reverse the order of the results and the hypothesis in their works, i.e. we construct averaged versions of Fokker-Planck equations using the above structure of non-abelian zeta zeros. This then leads to an infinite dimensional Hilbert vector bundle with smooth sections parametrized by non-abelian zeta zeros. We conjectures that the above structure of Fokker-Planck equation comes naturally from an ‘essential projectively flat connection’ of the infinite dimensional Hilbert bundle and the above family of smooth sections are its ‘essential pro-flat sections’.

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1 Non-abelian Zeta Functions

1.1 Definition

By definition, a *rank n lattice* Λ is a discrete subgroup of full rank in an n -dimensional Euclidean space. It is called *semi-stable* if for any sub lattice Λ_1 of Λ , we have

$$\text{vol}(\Lambda_1)^n \geq \text{vol}(\Lambda)^{n_1}. \quad (1)$$

Here, n_1 denotes the rank of Λ_1 and $\text{vol}(\Lambda)$ and $\text{vol}(\Lambda_1)$ denote the (co)volumes of Λ and Λ_1 , respectively.

Denote by $\mathcal{M}_{\mathbb{Q},n}$ be the moduli space of semi-stable lattices of rank n , parametrized the isometric classes of all rank n semi-stable lattices. It is well known that (see [5])

- (1) There is a natural smooth fibration $\text{vol} : \mathcal{M}_{\mathbb{Q},n} \rightarrow \mathbb{R}_{>0}$, sending the isometric classes $[\Lambda]$ of semi-stable lattices Λ to their volumes $\text{vol}(\Lambda)$.
- (2) For $t \in \mathbb{R}_{>0}$, denote the fiber of vol over t by $\mathcal{M}_{\mathbb{Q},n}[t]$. Obviously, $\mathcal{M}_{\mathbb{Q},n}[t]$ coincides with the moduli spaces of semi-stable lattices of rank n and volume t . Moreover, as a topological space, $\mathcal{M}_{\mathbb{Q},n}[t]$ is compact.
- (3) The natural embeddings $\mathcal{M}_{\mathbb{Q},n} \hookrightarrow \text{GL}_n(\mathbb{Z}) \backslash \text{GL}_n(\mathbb{R})$ induces compatible embeddings $\mathcal{M}_{\mathbb{Q},n}[t] \hookrightarrow \text{SL}_n(\mathbb{Z}) \backslash \text{SL}_n(\mathbb{R})$. Thus, induced from the (Tamagawa) measures on the quotient spaces, we obtain natural volume forms $d\mu(\Lambda)$ and $d\mu_t(\Lambda)$ on $\mathcal{M}_{\mathbb{Q},n}$ and $\mathcal{M}_{\mathbb{Q},n}[t]$, respectively. Furthermore, we have

$$d\mu(\Lambda) = d\mu_0(\Lambda) \times \frac{dt}{t}. \quad (2)$$

With all these, we are ready to introduce the following

Definition 1 ([5]). *The rank n non-abelian zeta function $\widehat{\zeta}_{\mathbb{Q},n}(s)$ of the field \mathbb{Q} of rationals is defined by the integration*

$$\widehat{\zeta}_{\mathbb{Q},n}(s) := \int_{\mathcal{M}_{\mathbb{Q},n}} (\theta_{\Lambda} - 1) \cdot \text{vol}(\Lambda)^s d\mu(\Lambda) \quad \Re(s) > 1 \quad (3)$$

Here, as usual, $\theta_{\Lambda} := \sum_{x \in \Lambda} e^{-\pi|x|^2}$ denotes the theta series of the lattice Λ .

Obviously, when $n = 1$, $\widehat{\zeta}_{\mathbb{Q},n}(s)$ coincides with the (completed) Riemann zeta function.

1.2 Riemann Hypothesis

The non-abelian zeta function $\widehat{\zeta}_{\mathbb{Q},n}(s)$ satisfies the Riemann properties for zeta functions. To explain this, we begin with the structural formula for non-abelian zeta function $\widehat{\zeta}_{\mathbb{Q},n}(s)$

$$s(s-1)\widehat{\zeta}_{\mathbb{Q},n}(s) = s(s-1)(I(s) + I(1-s)) + \text{vol}(\mathcal{M}_{\mathbb{Q},n}[1]) \quad (4)$$

where

$$I(s) := \int_{\mathcal{M}_{\mathbb{Q},n}[\leq 1]} \left(e^{h^0(\mathbb{Q},\Lambda)} - 1 \right) \cdot \text{vol}(\Lambda)^s d\mu(\Lambda), \quad (5)$$

and $h^i(\mathbb{Q}, \Lambda)$ ($i = 0, 1$) is the i -th arithmetic cohomology of the lattice Λ characterized by

$$h^0(\mathbb{Q}, \Lambda) = \log \left(\sum_{x \in \Lambda} e^{-\pi|x|_{\Lambda}^2} \right). \quad (6)$$

Since

$$\begin{aligned} \mathcal{M}_{\mathbb{Q},n}[t] &\simeq \mathcal{M}_{\mathbb{Q},n}[1] \\ \Lambda &\mapsto t^{1/n}\Lambda \end{aligned} \quad (7)$$

we have

$$\begin{aligned} I(s) &= \int_0^1 t^s \frac{dt}{t} \int_{\mathcal{M}_{\mathbb{Q},n}[1]} \left(e^{h^0(\mathbb{Q}, \frac{1}{t^{1/n}}\Lambda)} - 1 \right) d\mu(\Lambda) \\ &= \int_1^\infty t^{-s} \frac{dt}{t} \int_{\mathcal{M}_{\mathbb{Q},n}[1]} \left(e^{h^0(\mathbb{Q}, t^{1/n}\Lambda)} - 1 \right) d\mu(\Lambda) \\ &= \int_1^\infty t^{-s} \frac{dt}{t} \int_{\mathcal{M}_{\mathbb{Q},n}[1]} \left(\sum_{x \in \Lambda \setminus \{0\}} e^{-\pi t^{2/n}|x|_{\Lambda}^2} \right) d\mu(\Lambda) \\ &= \frac{n}{2} \int_1^\infty t^{-\frac{n}{2}s} \frac{dt}{t} \int_{\mathcal{M}_{\mathbb{Q},n}[1]} \left(\sum_{x \in \Lambda \setminus \{0\}} e^{-\pi t|x|_{\Lambda}^2} \right) d\mu(\Lambda). \end{aligned} \quad (8)$$

Put this back to (5), we obtain

$$\begin{aligned} &2s(s-1)\widehat{\zeta}_{\mathbb{Q},n}(s) - 2\text{vol}(\mathcal{M}_{\mathbb{Q},n}[1]) \\ &= ns(s-1) \int_1^\infty \left(t^{-\frac{n}{2}s} + t^{-\frac{n}{2}(1-s)} \right) \frac{dt}{t} \left(\int_{\mathcal{M}_{\mathbb{Q},n}[1]} \left(\sum_{x \in \Lambda \setminus \{0\}} e^{-\pi t|x|_{\Lambda}^2} \right) d\mu(\Lambda) \right) \\ &= ns(s-1) \int_1^\infty \left(t^{-\frac{n}{2}s} + t^{-\frac{n}{2}(1-s)} \right) \frac{dt}{t} \left(\int_{\mathcal{M}_{\mathbb{Q},n}[1]} \left(\theta_{\Lambda}(t) - 1 \right) d\mu(\Lambda) \right). \end{aligned} \quad (9)$$

Hence, at least formally, we obtain the *functional equation* for $\widehat{\zeta}_{\mathbb{Q},n}(s)$, i.e.

$$\widehat{\zeta}_{\mathbb{Q},n}(1-s) = \widehat{\zeta}_{\mathbb{Q},n}(s). \quad (10)$$

Moreover, by the *arithmetic vanishing theorem* on $h^1(\mathbb{Q}, \Lambda)$ and arithmetic duality between $h^0(\mathbb{Q}, \Lambda)$ and $h^1(\mathbb{Q}, \Lambda^\vee)$, we can show that $I(s)$ is a holomorphic function in s . Therefore, the non-abelian zeta function $\widehat{\zeta}_{\mathbb{Q},n}(s)$ is well-defined

and admits a unique meromorphic continuation to the whole s -plane, which admits only two singularities, i.e. two simple at $s = 0, 1$ with the residue at $s = 1$ given by the volume of the compact moduli space $\mathcal{M}_{\mathbb{Q},n}[1]$. As a by-product, (10) is justified. Above all these, much more surprisingly, we have the following result on the Riemann hypothesis of $\widehat{\zeta}_{\mathbb{Q},n}(s)$.

Theorem 2 (Theorem 15.4 of [5]). *Assume $n \geq 2$. Then all but finitely many zeros of $\widehat{\zeta}_{\mathbb{Q},n}(s)$ lie on the line $\Re(s) = \frac{1}{2}$.*

For the basic facts of $\widehat{\zeta}_{\mathbb{Q},n}(s)$, particularly, the proof of this theorem on weak Riemann hypothesis for $\widehat{\zeta}_{\mathbb{Q},n}(s)$, please refer to [5], which studies new yet genuine zeta functions for reductive groups over number fields systematically.

2 Heat Equations and Resolvent Formulas

Recall that a lattice Λ consists of a \mathbb{Z} -module P of rank n in the n -dimensional \mathbb{R} -vector space \mathbb{R}^n and an Euclidean metric H on \mathbb{R}^n . Let $H = (h^{\alpha\beta})$ be the positive definite symmetric matrix associated to the rank n lattice Λ , $x = (x^1, x^2, \dots, x^n)$ and set

$$\theta_{\Lambda}(x, t) := \sum_{\lambda \in \Lambda} e^{-\pi t |\lambda|_{\Lambda}^2} e^{2\pi i \langle \lambda, x \rangle} \quad (11)$$

be the theta function associated to the lattice $\Lambda = (P, H)$. Here, in the equation above, we have set

$$|(\lambda_1, \dots, \lambda_n)|_{\Lambda}^2 = \sum_{\alpha, \beta=1}^n h^{\alpha\beta} \lambda_{\alpha} \lambda_{\beta} \quad \text{and} \quad \langle \lambda, x \rangle = \sum_{\alpha=1}^n \lambda_{\alpha} x^{\alpha} \quad (12)$$

As usual, we define the *dual lattice* Λ^{\vee} of the lattice Λ and its induced torus $\mathbb{T}_{\Lambda^{\vee}}^n$ by

$$\Lambda^{\vee} := \{x \in \mathbb{R}^n : \langle \lambda, x \rangle \in \mathbb{Z} \forall \lambda \in \Lambda\} \quad \text{and} \quad \mathbb{T}_{\Lambda^{\vee}}^n := \mathbb{R}^n / \Lambda^{\vee}, \quad (13)$$

respectively. Then $\theta(x, t)$ may be viewed as a genuine function on the torus $\mathbb{T}_{\Lambda^{\vee}}^n$. Since

$$\frac{\partial}{\partial x^{\alpha}} \theta_{\Lambda}(x, t) = (2\pi i) \sum_{\lambda \in \Lambda} e^{-\pi t |\lambda|_{\Lambda}^2} e^{2\pi i \langle \lambda, x \rangle} \lambda_{\alpha}, \quad (14)$$

we have

$$\frac{\partial^2}{\partial x^{\beta} \partial x^{\alpha}} \theta_{\Lambda}(x, t) = 4\pi^2 \sum_{\lambda \in \Lambda} e^{-\pi t |\lambda|_{\Lambda}^2} e^{2\pi i \langle \lambda, x \rangle} \lambda_{\alpha} \lambda_{\beta} \quad (15)$$

Therefore,

$$\begin{aligned} \sum_{\alpha, \beta=1}^n h^{\alpha\beta} \frac{\partial^2}{\partial x^{\beta} \partial x^{\alpha}} \theta_{\Lambda}(x, t) &= 4\pi^2 \sum_{\lambda \in \Lambda} e^{-\pi t |\lambda|_{\Lambda}^2} e^{2\pi i \langle \lambda, x \rangle} \sum_{\alpha, \beta=1}^n h^{\alpha\beta} \lambda_{\alpha} \lambda_{\beta} \\ &= 4\pi^2 \sum_{\lambda \in \Lambda} e^{-\pi t |\lambda|_{\Lambda}^2} e^{2\pi i \langle \lambda, x \rangle} |\lambda|_{\Lambda}^2 \end{aligned} \quad (16)$$

On the other hand, since

$$\frac{\partial}{\partial t} \theta_{\Lambda}(x, t) = -\pi \sum_{\lambda \in \Lambda} e^{-\pi t |\lambda|_{\Lambda}^2} e^{2\pi i \langle \lambda, x \rangle} |\lambda|_{\Lambda}^2 \quad (17)$$

all these then verify the following well-known

Lemma 3 (Heat Equation). *On the torus $\mathbb{T}_{\Lambda^{\vee}}^n$ associated to the dual lattice Λ^{\vee} of the lattice Λ , the theta function $\theta_{\Lambda}(x, t)$ of Λ is a fundamental solution of the following heat equation*

$$\begin{cases} \frac{1}{4\pi} \Delta_{\Lambda^{\vee}, x} (\theta_{\Lambda}(x, t)) = \frac{\partial}{\partial t} (\theta_{\Lambda}(x, t)), \\ \theta_{\Lambda}(x, 0) = \sum_{\lambda \in \Lambda} e^{2\pi i \langle \lambda, x \rangle}. \end{cases} \quad (18)$$

Here $\Delta_{\Lambda^{\vee}, x}$ denotes the Laplacian operator

$$\Delta_{\Lambda^{\vee}, x} := \sum_{\alpha, \beta=1}^n h^{\alpha\beta} \partial_{\alpha} \partial_{\beta} \quad (19)$$

with $\partial_{\alpha} := \frac{\partial}{\partial x^{\alpha}}$, $\partial_{\beta} := \frac{\partial}{\partial x^{\beta}}$.

Remark 1. Even the natural Laplace operator on the tangent bundle of $\mathbb{T}_{\Lambda^{\vee}}^n$ and hence on $\mathbb{T}_{\Lambda^{\vee}}^n$ is given by

$$\Delta_{\Lambda^{\vee}} := \sum_{i, \alpha=1}^n h_{\alpha\beta} \partial_{\alpha} \partial_{\beta} \quad (20)$$

with $(h_{\alpha\beta})$ the inverse matrix of $H = (h^{\alpha\beta})$, for our purpose, we decide to use the 'dual' Laplace operator $\Delta_{\Lambda^{\vee}, x}$. This operator $\Delta_{\Lambda^{\vee}, x}$, viewed as the Laplacian on the bundle of differential forms of $\mathbb{T}_{\Lambda^{\vee}}^n$, is also quite natural. Hope that this would not lead to any notational confusion.

To facilitate our further discussion, set now $t := \exp(T)$, $x := \exp(T/2) X$,

$$\Delta_{\Lambda^{\vee}, X} := \sum_{\alpha, \beta} h^{\alpha\beta} \frac{\partial^2}{\partial X^{\alpha} \partial X^{\beta}}, \quad \left\langle X, \frac{\partial}{\partial X} \right\rangle := \sum_{\alpha=1}^n X_{\alpha} \frac{\partial}{\partial X_{\alpha}}, \quad (21)$$

and

$$\Theta_{\Lambda}(X, T) := \theta_{\Lambda}(X \exp(T/2), \exp T) = \sum_{\lambda \in \Lambda} e^{-\pi e^T |\lambda|_{\Lambda}^2} e^{2\pi i \langle \lambda, X \rangle e^{T/2}}. \quad (22)$$

Then,

$$\begin{aligned} \Delta_{\Lambda^{\vee}, x} &= \sum_{\alpha, \beta} h^{\alpha\beta} \frac{\partial^2}{\partial (X^{\alpha} \exp(T/2)) \partial (X^{\beta} \exp(T/2))} = e^{-T} \Delta_{\Lambda^{\vee}, X}, \\ \frac{1}{4\pi} \Delta_{\Lambda^{\vee}, X} (\Theta_{\Lambda}(X, T)) &= -\pi e^T \sum_{\lambda \in \Lambda} |\lambda|_{\Lambda}^2 e^{-\pi e^{T/2} |\lambda|_{\Lambda}^2} e^{2\pi i \langle \lambda, X \rangle e^{T/2}}. \end{aligned} \quad (23)$$

On the other hand, since

$$\begin{aligned}
\frac{\partial}{\partial T} \left(\Theta_\Lambda(X, T) \right) &= \frac{\partial}{\partial T} \left(\sum_{\lambda \in \Lambda} e^{-\pi e^T |\lambda|_\Lambda^2} e^{2\pi i \langle \lambda, X \rangle} e^{T/2} \right) \\
&= \sum_{\lambda \in \Lambda} e^{\pi e^{T/2} |\lambda|_\Lambda^2} e^{2\pi i \langle \lambda, X \rangle} e^{T/2} \left(-\pi e^T |\lambda|_\Lambda^2 + \pi i \langle \lambda, X \rangle e^{T/2} \right) \\
&= \frac{1}{4\pi} \Delta_{\Lambda^\vee, X} \left(\Theta_\Lambda(X, T) \right) + \pi i e^{T/2} \sum_{\lambda \in \Lambda} \langle \lambda, X \rangle e^{\pi e^{T/2} |\lambda|_\Lambda^2} e^{2\pi i \langle \lambda, X \rangle} e^{T/2},
\end{aligned} \tag{24}$$

and

$$\begin{aligned}
\left\langle X, \frac{\partial}{\partial X} \right\rangle \left(\Theta_\Lambda(X, T) \right) &= \sum_{\beta=1}^n X^\beta \left(\sum_{\lambda \in \Lambda} e^{\pi e^T |\lambda|_\Lambda^2} e^{2\pi i \langle \lambda, X \rangle} e^{T/2} \left(2\pi i \lambda_\beta e^{T/2} \right) \right) \\
&= 2\pi i e^{T/2} \sum_{\lambda \in \Lambda} \langle \lambda, X \rangle e^{\pi e^{T/2} |\lambda|_\Lambda^2} e^{2\pi i \langle \lambda, X \rangle} e^{T/2}
\end{aligned} \tag{25}$$

we have proved the following theorem by using the obvious relation

$$\Theta_\Lambda(X, 0) = \sum_{\lambda \in \Lambda} e^{\pi |\lambda|_\Lambda^2} e^{2\pi i \langle \lambda, X \rangle} = \theta_\Lambda(\lambda, 1). \tag{26}$$

Lemma 4. *On the dual torus $\mathbb{T}_{\Lambda^\vee} \simeq \mathbb{R}^n / (e^{-T/2} \Lambda^\vee)$, we have*

$$\begin{cases} \left(\frac{\partial}{\partial T} - \Omega_{\Lambda^\vee, X} \right) \left(\Theta_\Lambda(X, T) \right) = 0, \\ \Theta_\Lambda(X, 0) = \theta_\Lambda(x, 1). \end{cases} \tag{27}$$

Here, $\Omega_{\Lambda^\vee, X}$ denotes the second order differential operator defined by

$$\Omega_{\Lambda^\vee, X} := \frac{1}{4\pi} \Delta_{\Lambda^\vee, X} + \frac{1}{2} \left\langle X, \frac{\partial}{\partial X} \right\rangle. \tag{28}$$

From (28), we see that the second differential operator $\Omega_{\Lambda^\vee, X}$ is elliptic as well.

3 Relation with Non-Abelian Zeta Functions

Now we are ready to explain why a resolvent version of the discussion above can be naturally connected to the rank n zeta function. To start with, by (9), in terms of variables X and T , we have

$$\begin{aligned}
&2s(s-1) \widehat{\zeta}_{\mathbb{Q}, n}(s) - 2 \text{vol}(\mathcal{M}_{\mathbb{Q}, n}[1]) \\
&= ns(s-1) \int_{\mathcal{M}_{\mathbb{Q}, n}[1]} \left(\int_0^\infty \left(e^{\frac{nT}{2}(1-s)} + e^{\frac{nT}{2}s} \right) \left(\Theta_\Lambda(0, T) - 1 \right) dT \right) d\mu(\Lambda)
\end{aligned} \tag{29}$$

Obviously, the integrant for the outer integration over $\mathcal{M}_{\mathbb{Q}, n}[1]$, namely

$$\int_0^\infty \left(e^{\frac{nT}{2}(1-s)} + e^{\frac{nT}{2}s} \right) \left(\Theta_\Lambda(0, T) - 1 \right) dT, \tag{30}$$

involves two Laplace transforms

$$\int_0^\infty e^{\frac{nT}{2}(1-s)} (\Theta_\Lambda(0, T) - 1) dT \quad \text{and} \quad \int_0^\infty e^{\frac{nT}{2}s} (\Theta_\Lambda(0, T) - 1) dT, \quad (31)$$

hence there is a contraction semi-group and hence a Feller process induces by the heat equation

$$\Omega_{\Lambda^\vee, X} (\Theta_\Lambda(X, T) - 1) = \frac{\partial}{\partial T} (\Theta_\Lambda(X, T) - 1). \quad (32)$$

Consequently, we may apply the general resolvent formula (see e.g. p.316 of [3]) to obtain

$$\int_0^\infty e^{\kappa T} (\Theta_\Lambda(X, T) - 1) dT = R(\kappa, \Omega_{\Lambda^\vee, X}) (\Theta_\Lambda(X, T) - 1). \quad (33)$$

Here, $R(\kappa, \Omega_{\Lambda^\vee, X})$ denotes the resolvent associated to $(-\kappa - \Omega_{\Lambda^\vee, X})^{-1}$. Therefore, we obtain the following

Corollary 5. *For the rank n non-abelian zeta function $\widehat{\zeta}_{\mathbb{Q}, n}(s)$, we have*

$$\begin{aligned} & \frac{2}{n} \left(\widehat{\zeta}_{\mathbb{Q}, n}(s) - \frac{1}{s(s-1)} \text{vol}(\mathcal{M}_{\mathbb{Q}, n}[1]) \right) \\ &= \int_{\mathcal{M}_{\mathbb{Q}, n}[1]} \left(\left(R\left(-\frac{ns}{2}, \Omega_{\Lambda^\vee, X}\right) + R\left(-\frac{n(1-s)}{2}, \Omega_{\Lambda^\vee, X}\right) \right) (\Theta_\Lambda(X, 0) - 1) \right) \Big|_{X=0} d\mu(\Lambda). \end{aligned} \quad (34)$$

We end this section by pointing out what are the global structures involved in (34). Over the moduli space $\mathcal{M}_{\mathbb{Q}, n}$ of semi-stable lattices of rank n , naturally associated is the universal family $\pi : \mathcal{T}_{\mathbb{Q}, n}^\vee \rightarrow \mathcal{M}_{\mathbb{Q}, n}$, characterized by the property that, for each point $[\Lambda] \in \mathcal{M}_{\mathbb{Q}, n}$, the fiber $\pi^{-1}([\Lambda])$ is simply the torus $\mathbb{T}_{\Lambda^\vee} := \mathbb{R}^n / \Lambda^\vee$ associated to the dual lattice Λ^\vee of Λ . Therefore, on the right hand side of (34), namely in

$$\int_{\mathcal{M}_{\mathbb{Q}, n}[1]} \left(\left(R\left(-\frac{ns}{2}, \Omega_{\Lambda^\vee, X}\right) + R\left(-\frac{n(1-s)}{2}, \Omega_{\Lambda^\vee, X}\right) \right) (\Theta_\Lambda(X, 0) - 1) \right) \Big|_{X=0} d\mu(\Lambda), \quad (35)$$

the operators $\Omega_{\Lambda^\vee, X}$ are second order elliptic differential operators on the fibers $\mathbb{T}_{\Lambda^\vee}$ of π , the so-called vertical direction, while the integration is on the horizontal direction of the base moduli space $\mathcal{M}_{\mathbb{Q}, n}[1]$ of the semi-stable lattices Λ of rank n . Therefore, the whole right hand side is nothing but a horizontal average on the base space of π for the 'action' of the resolvents on the theta series $\Theta_\Lambda(X, 0)$ along the vertical fiber direction of π .

4 Non-Abelian Zeta Zeros as Eigenvalues

In the integration of (35), the integrand is taken along the special section $X = 0$ in of the fibration π . But this is artificial: To recover the vertical direction, it is enough to assume that X is arbitrary. As to be expect, for this to work, a price has to be paid with a restriction in continuous parameter s , so as to obtain certain initial condition to stabilize the solutions of our averaged differential equation. It is for this purpose that we have to focus on the zeros of $\widehat{\zeta}_{\mathbb{Q}, n}(s)$.

By the weak Riemann hypothesis established in Theorem 2, when $n \geq 2$, all but finitely many zeros of $\widehat{\zeta}_{\mathbb{Q},n}(s)$ lie on the central line $\Re(s) = \frac{1}{2}$. Let then $\rho = \frac{1}{2} + i\gamma \in \frac{1}{2} + i\mathbb{R}$ be a zero of $\widehat{\zeta}_{\mathbb{Q},n}(s)$. Then (34) becomes

$$\begin{aligned} & \frac{2}{n} \left(\widehat{\zeta}_{\mathbb{Q},n} \left(\frac{1}{2} + i\gamma \right) - \text{vol}(\mathcal{M}_{\mathbb{Q},n}[1]) \frac{1}{1/4 + \gamma^2} \right) \\ &= \int_{\mathcal{M}_{\mathbb{Q},n}[1]} \left(\left(R \left(-\frac{n}{4} - \frac{n}{2} \gamma i, \Omega_{\Lambda^\vee, X} \right) + R \left(-\frac{n}{4} + \frac{n}{2} \gamma i, \Omega_{\Lambda^\vee, X} \right) \right) (\Theta_\Lambda(X, 0) - 1) \right) \Big|_{X=0} d\mu(\Lambda) \\ &= \int_{\mathcal{M}_{\mathbb{Q},n}[1]} d\mu(\Lambda) \int_0^\infty \left(e^{(\frac{n}{4} + \frac{n}{2} i\gamma)T} + e^{(\frac{n}{4} - \frac{n}{2} i\gamma)T} \right) \sum_{\lambda \in \Lambda \setminus \{0\}} e^{-\pi e^T |\lambda|_\Lambda^2} e^{2\pi i \langle \lambda, X \rangle} e^{T/2} \Big|_{X=0} dT. \end{aligned} \quad (36)$$

Accordingly, we introduce a family of functions $\widehat{\zeta}_{\mathbb{Q},n}(X, \frac{1}{2} + i\gamma)$ of X parametrized by γ as follows

$$\begin{aligned} & \frac{2}{n} \left(\widehat{\zeta}_{\mathbb{Q},n} \left(X, \frac{1}{2} + i\gamma \right) - \text{vol}(\mathcal{M}_{\mathbb{Q},n}[1]) \frac{1}{1/4 + \gamma^2} \right) \\ &:= \int_{\mathcal{M}_{\mathbb{Q},n}[1]} d\mu(\Lambda) \int_0^\infty \left(e^{(\frac{n}{4} + \frac{n}{2} i\gamma)T} + e^{(\frac{n}{4} - \frac{n}{2} i\gamma)T} \right) \sum_{\lambda \in \Lambda \setminus \{0\}} e^{-\pi e^T |\lambda|_\Lambda^2} e^{2\pi i \langle \lambda, X \rangle} e^{T/2} dT \\ &= \int_{\mathcal{M}_{\mathbb{Q},n}[1]} \left(\left(R \left(-\frac{n}{4} - \frac{n}{2} \gamma i, \Omega_{\Lambda^\vee, X} \right) + R \left(-\frac{n}{4} + \frac{n}{2} \gamma i, \Omega_{\Lambda^\vee, X} \right) \right) (\Theta_\Lambda(X, 0) - 1) \right) d\mu(\Lambda) \\ &= 4 \int_{\mathcal{M}_{\mathbb{Q},n}[1]} \left(R \left(-(n\gamma)^2, \left(2\Omega_{\Lambda^\vee, X} + \frac{n}{2} \right)^2 \right) \left(\left(2\Omega_{\Lambda^\vee, X} + \frac{n}{2} \right) (\Theta_\Lambda(X, 0) - 1) \right) \right) d\mu(\Lambda) \end{aligned} \quad (37)$$

Here, in the last step, we have used the following elementary identity

$$(\Omega - \alpha)^{-1} + (\Omega - \beta)^{-1} = ((\Omega - \alpha)(\Omega - \beta))^{-1} (2\Omega - (\alpha + \beta)). \quad (38)$$

To simplify our notations, set

$$\begin{aligned} \Phi_{\Lambda^\vee}^{(1)}(X, \gamma) &:= \frac{n}{2} \int_0^\infty \left(e^{(\frac{n}{4} + \frac{n}{2} i\gamma)T} + e^{(\frac{n}{4} - \frac{n}{2} i\gamma)T} \right) (\Theta_\Lambda(X, T) - 1) dT, \\ \kappa_{\Lambda^\vee}^{(1)}(X) &:= 2n \left(2\Omega_{\Lambda^\vee, X} + \frac{n}{2} \right) (\Theta_\Lambda(X, 0) - 1). \end{aligned} \quad (39)$$

Since X is a flat section of the fibration π , the relation (37) implies that

$$\begin{aligned} & \left(\left(2\Omega_{\Lambda^\vee, X} + \frac{n}{2} \right)^2 + (n\gamma)^2 \right) \left(\widehat{\zeta}_{\mathbb{Q},n} \left(X, \frac{1}{2} + i\gamma \right) - \frac{\text{vol}(\mathcal{M}_{\mathbb{Q},n}[1])}{1/4 + \gamma^2} \right) \\ &= \int_{\mathcal{M}_{\mathbb{Q},n}[1]} \left(\left(\left(2\Omega_{\Lambda^\vee, X} + \frac{n}{2} \right)^2 + (n\gamma)^2 \right) \Phi_{\Lambda^\vee}^{(1)}(X, \gamma) \right) d\mu(\Lambda) \\ &= \int_{\mathcal{M}_{\mathbb{Q},n}[1]} \kappa_{\Lambda^\vee}^{(1)}(X) d\mu(\Lambda) \end{aligned} \quad (40)$$

To obtain an initial condition, note that

$$\widehat{\zeta}_{\mathbb{Q},n} \left(0, \frac{1}{2} + i\gamma \right) = 0, \quad (41)$$

we hence arrive at the normalizations

$$\begin{aligned}
& \left(\left(2\Omega_{\Lambda^\vee, X} + \frac{n}{2} \right)^2 + (n\gamma)^2 \right) \left(\widehat{\zeta}_{\mathbb{Q}, n} \left(X, \frac{1}{2} + i\gamma \right) - \frac{\text{vol}(\mathcal{M}_{\mathbb{Q}, n}[1])}{1/4 + \gamma^2} (1 - \exp(-\pi|X|^2)) \right) \\
&= \int_{\mathcal{M}_{\mathbb{Q}, n}[1]} \left(\left(\left(2\Omega_{\Lambda^\vee, X} + \frac{n}{2} \right)^2 + (n\gamma)^2 \right) \Phi_{\Lambda^\vee}(X, \gamma) \right) d\mu(\Lambda) \\
&= \int_{\mathcal{M}_{\mathbb{Q}, n}[1]} \kappa_{\Lambda^\vee}(X) d\mu(\Lambda).
\end{aligned} \tag{42}$$

where, as to be verified with a long but tedious calculation, the functions $\Phi_{\Lambda^\vee}(X, \gamma)$ and $\kappa_{\Lambda^\vee}(X)$ are given by

$$\begin{aligned}
\Phi_{\Lambda^\vee}(X, \gamma) &= \Phi_{\Lambda^\vee}^{(1)}(X, \gamma) + \frac{1}{1/4 + \gamma^2} \exp(-\pi|X|^2), \\
\kappa_{\Lambda^\vee}(X) &= \kappa_{\Lambda^\vee}^{(1)}(X) + \frac{1}{1/4 + \gamma^2} \left(\left(2\Omega_{\Lambda^\vee, X} + \frac{n}{2} \right)^2 + (n\gamma)^2 \right) \exp(-\pi|X|^2)
\end{aligned} \tag{43}$$

In this way, we complete a proof of the following

Theorem 6. *Assume that $\frac{1}{2} + i\gamma \in \frac{1}{2} + i\mathbb{R}$ is a zero of the non-abelian zeta function $\widehat{\zeta}_{\mathbb{Q}, n}(s)$. Then the function $\Phi_{\Lambda^\vee}(X, \gamma)$ is a solution of the following ‘average’ differential equation*

$$\int_{\mathcal{M}_{\mathbb{Q}, n}[1]} \left(\left(\left(2\Omega_{\Lambda^\vee, X} + \frac{n}{2} \right)^2 + (n\gamma)^2 \right) \Phi_{\Lambda^\vee}(X, \gamma) - \kappa_{\Lambda^\vee}(X) \right) d\mu(\Lambda) = 0 \tag{44}$$

and satisfies the following initial conditions

$$\begin{aligned}
(1) \quad & \int_{\mathcal{M}_{\mathbb{Q}, n}[1]} \Phi_{\Lambda^\vee}(0, \gamma) d\mu(\Lambda) = 0, \\
(2) \quad & \int_{\mathcal{M}_{\mathbb{Q}, n}[1]} \frac{\partial}{\partial X_i} \Phi_{\Lambda^\vee}(0, \gamma) d\mu(\Lambda) = \int_{\mathcal{M}_{\mathbb{Q}, n}[1]} \frac{\partial^3}{\partial X_i^3} \Phi_{\Lambda^\vee}(0, \gamma) d\mu(\Lambda) = 0; \\
(3) \quad & \lim_{|X| \rightarrow \infty} \left| \int_{\mathcal{M}_{\mathbb{Q}, n}[1]} \frac{\partial^j}{\partial X_i^j} \Phi_{\Lambda^\vee}(X, \gamma) d\mu(\Lambda) \right| = O\left(\frac{1}{|X|}\right) \text{ for } 0 \leq j \leq 4.
\end{aligned}$$

Proof. Indeed, (44) is a direct consequence of (36) and (42) as mentioned above. Since $\widehat{\zeta}_{\mathbb{Q}, n}(1/2 + i\gamma) = 0$, resp. $\Phi_{\Lambda^\vee}(X, \gamma) = \Phi_{\Lambda^\vee}(-X, \gamma)$, we have (1) and (2). In the same line, since the involvement of X in $\Phi_{\Lambda^\vee}(X, \gamma)$ is rather simple, we have (3) as well by direct calculations. \square

We may rewrite the integrand in (44) more uniformly in the language of perturbations. To explain this, we first give some functional analysis preparations.

Suppose there would be a family of Hilbert spaces $\mathcal{H}_{\Lambda, 0}$ of functions defined on the fibers $\mathbb{T}_{e^{-T/2}\Lambda^\vee}$ of $\pi : \mathcal{T}_{\mathbb{Q}, n} \rightarrow \mathcal{M}_{\mathbb{Q}, n}[1]$ with associated inner products (\cdot, \cdot) , containing both $\Phi_{\Lambda^\vee}(X, \gamma)$ and $\kappa_{\Lambda^\vee}(X)$ for all the γ 's. Denote the real and the imaginary part of $\Phi_{\Lambda^\vee}(X, \gamma)$ by $\alpha_{\Lambda, \gamma}(X)$ and $\beta_{\Lambda, \gamma}(X)$. Then

$$\Phi_{\Lambda^\vee}(X, \gamma) = \alpha_{\Lambda, \gamma}(X) + i\beta_{\Lambda, \gamma}(X) \quad \text{and} \quad \Phi_{\Lambda^\vee}(X, \bar{\gamma}) = \alpha_{\Lambda, \gamma}(X) - i\beta_{\Lambda, \gamma}(X). \tag{45}$$

In the spaces $\mathcal{H}_{\Lambda, 0}$, chose a family of function $\varphi_\Lambda(X)$ such that

$$\langle \beta_{\Lambda, \gamma}(X), \varphi_\Lambda(X) \rangle = 0 \quad \text{and} \quad \langle \alpha_{\Lambda, \gamma}(X), \varphi_\Lambda(X) \rangle \neq 0. \tag{46}$$

We then could define a family of projection operators

$$\begin{aligned} P_{\Lambda,\gamma} : \mathcal{H}_{\Lambda,0} &\rightarrow \langle \kappa_{\Lambda^\vee}(X) \rangle, \\ f &\mapsto -\frac{\langle f, \varphi_\Lambda(X) \rangle}{\langle \alpha_{\Lambda,\gamma}(X), \varphi_\Lambda(X) \rangle} \cdot \kappa_{\Lambda^\vee}(X) \end{aligned} \quad (47)$$

Certainly, $P_{\Lambda,\gamma}$ depends on γ and

$$P_{\Lambda,\bar{\gamma}} = P_{\Lambda,\gamma}. \quad (48)$$

Proposition 7. *Assume $\frac{1}{2} + \gamma$ is a zero of $\widehat{\zeta}_{\mathbb{Q},n}(s)$ on the central line. Then the average on $\mathcal{M}_{\mathbb{Q},n}[1]$ of the associated micro perturbations of (44) becomes stable in the sense that*

$$\int_{\mathcal{M}_{\mathbb{Q},n}[1]} \left(\left(\left(2\Omega_{\Lambda^\vee,X} + \frac{n}{2} \right)^2 + P_{\Lambda,\gamma} + (n\gamma)^2 \right) \Phi_{\Lambda^\vee}(X, \gamma) \right) d\mu(\Lambda) = 0. \quad (49)$$

When, $n \geq 2$, all but finitely many γ are real. Hence, by taking an average over $\mathcal{M}_{\mathbb{Q},n}[1]$, the number $-(n\gamma)^2$ is a sort of common real eigenvalue of the perturbed operator $\left(2\Omega_{\Lambda^\vee,X} + \frac{n}{2} \right)^2 + P_{\Lambda,\gamma}$ on a fixed family of Hilbert spaces $\mathcal{H}_{\Lambda,0}$, or better, a bundle of Hilbert spaces of π .

5 Fokker-Planck Equations and Beyond

5.1 Fokker-Planck Equations

Consider the basic differential equation stated in Theorem 6

$$\int_{\mathcal{M}_{\mathbb{Q},n}[1]} \left(\left(\left(2\Omega_{\Lambda^\vee,X} + \frac{n}{2} \right)^2 + (n\gamma)^2 \right) \Phi_{\Lambda^\vee}(X, \gamma) - \kappa_{\Lambda^\vee}(X) \right) d\mu(\Lambda) = 0. \quad (50)$$

Since all the functions involved define tempered distributions on the fibers $\mathbb{T}_{\Lambda^\vee}$ of the universal families $\pi : \mathcal{T}_{\mathbb{Q},n}^\vee \rightarrow \mathcal{M}_{\mathbb{Q},n}$ at $[\Lambda]$, we may take fiberwise Fourier transforms (for details, see e.g. §?? in the the appendices). Namely, along each fiber $\mathbb{T}_{\Lambda^\vee}$,

$$\widehat{f}(Y) := \mathcal{F}(f(X)) := \int_{-\infty}^{\infty} \dots \int_{-\infty}^{\infty} e^{-2\pi i \langle Y, X \rangle} f(X) dX \quad (51)$$

where dX is induced from a smooth family of volume forms on \mathbb{T}_Λ , compatible with the fibration structure of the universal family $\pi : \mathcal{T} \rightarrow \mathcal{M}_{\mathbb{Q},n}$ on $\mathcal{M}_{\mathbb{Q},n}$. Hence the following formulas hold tautologically

$$\widehat{D^a f} = (2\pi i x)^a \widehat{f} \quad \text{and} \quad (-2\pi i x)^a \widehat{f} = D^a \widehat{f} \quad (52)$$

where, for $a = (a_1, \dots, a_n)$, $|a| := \sum_{\alpha=1}^n a_\alpha$ and

$$x^a := \prod_{\alpha=1}^n x_i^{a_i} \quad \text{and} \quad \frac{\partial^{|a|}}{\partial x^a} := \frac{\partial^{|a|}}{\partial x_1^{a_1} \dots \partial x_n^{a_n}}. \quad (53)$$

Consequently, we have

$$(x^\alpha \partial_\alpha f)^\widehat{=} = \frac{1}{-2\pi i} \partial_\alpha (\partial_\alpha f)^\widehat{=} = \frac{2\pi i}{-2\pi i} \partial_\alpha (x^\alpha \widehat{f}) = -\partial_\alpha (x^\alpha \widehat{f}). \quad (54)$$

Therefore, ‘dual’ to (50), with a direct but tedious calculation, we conclude that

$$\begin{aligned}
& \int_{\mathcal{M}_{\mathbb{Q},n}^{[1]}} \left(\sum_{\alpha,\beta=1}^n \frac{\partial}{\partial Y_\alpha} \left(Y_\alpha \frac{\partial}{\partial Y_\beta} \left(Y_\beta \widehat{\Phi}_{\Lambda^\vee}(Y, \gamma) \right) \right) - n \sum_{\alpha=1}^n \frac{\partial}{\partial Y_\alpha} \left(Y_\alpha \widehat{\Phi}_{\Lambda^\vee}(Y, \gamma) \right) \right. \\
& + 2\pi \left(\sum_{\alpha,\beta,\tau=1}^n h^{\alpha\beta} Y_\alpha Y_\beta \frac{\partial}{\partial Y_\tau} \left(Y_\tau \widehat{\Phi}_{\Lambda^\vee}(Y, \gamma) \right) + \sum_{\alpha,\beta,\tau=1}^n h^{\alpha\beta} \frac{\partial}{\partial Y_\tau} (Y_\alpha Y_\beta Y_\tau \widehat{\Phi}_{\Lambda^\vee}(Y, \gamma)) \right) \\
& + 4\pi^2 \left(\sum_{\alpha,\beta=1}^n h^{\alpha\beta} Y_\alpha Y_\beta \right)^2 \widehat{\Phi}_{\Lambda^\vee}(Y, \gamma) - 2\pi n \sum_{\alpha,\beta=1}^n h^{\alpha\beta} Y_\alpha Y_\beta \widehat{\Phi}_{\Lambda^\vee}(Y, \gamma) \\
& \left. + n^2 \left(\frac{1}{4} + \gamma^2 \right) \widehat{\Phi}_{\Lambda^\vee}(Y, \gamma) - \widehat{\kappa_{\Lambda^\vee}}(Y) \right) d\mu = 0
\end{aligned} \tag{55}$$

To continue, we next give the following elementary lemma, which can be deduced from a long but tedious elementary calculation, using the following elementary relation repeatedly

$$Y_\alpha \frac{\partial}{\partial Y_\alpha} (A) = \frac{\partial}{\partial Y_\alpha} (Y_\alpha A) - A \tag{56}$$

where A is a C^1 function in Y .

Lemma 8. *The following identities hold.*

$$\begin{aligned}
& \sum_{\alpha,\beta=1}^n \frac{\partial}{\partial Y_\alpha} \left(Y_\alpha \frac{\partial}{\partial Y_\beta} \left(Y_\beta \widehat{\Phi}_{\Lambda^\vee}(Y, \gamma) \right) \right) \\
& = n^2 \widehat{\Phi}_{\Lambda^\vee}(Y, \gamma) + (2n+1) \sum_{\alpha=1}^n Y_\alpha \frac{\partial}{\partial Y_\alpha} \left(\widehat{\Phi}_{\Lambda^\vee}(Y, \gamma) \right) + \sum_{\alpha,\beta=1}^n Y_\alpha Y_\beta \frac{\partial^2}{\partial Y_\alpha \partial Y_\beta} \left(\widehat{\Phi}_{\Lambda^\vee}(Y, \gamma) \right), \\
& n \sum_{\alpha=1}^n \frac{\partial}{\partial Y_\alpha} \left(Y_\alpha \widehat{\Phi}_{\Lambda^\vee}(Y, \gamma) \right) = n^2 \widehat{\Phi}_{\Lambda^\vee}(Y, \gamma) + n \sum_{\alpha=1}^n Y_\alpha \frac{\partial}{\partial Y_\alpha} \left(\widehat{\Phi}_{\Lambda^\vee}(Y, \gamma) \right), \\
& \sum_{\alpha,\beta,\tau=1}^n h^{\alpha\beta} Y_\alpha Y_\beta \frac{\partial}{\partial Y_\tau} \left(Y_\tau \widehat{\Phi}_{\Lambda^\vee}(Y, \gamma) \right) = \sum_{\alpha,\beta,\tau=1}^n h^{\alpha\beta} Y_\alpha Y_\beta \left(1 + Y_\tau \frac{\partial}{\partial Y_\tau} \right) \left(\widehat{\Phi}_{\Lambda^\vee}(Y, \gamma) \right), \\
& \sum_{\alpha,\beta,\tau=1}^n h^{\alpha\beta} \frac{\partial}{\partial Y_\tau} (Y_\alpha Y_\beta Y_\tau \widehat{\Phi}_{\Lambda^\vee}(Y, \gamma)) = - \sum_{\alpha=1}^n h^{\alpha\alpha} Y_\alpha^2 \left(\widehat{\Phi}_{\Lambda^\vee}(Y, \gamma) \right) \\
& + (n+3) \sum_{\alpha,\beta=1}^n h^{\alpha\beta} Y_\alpha Y_\beta \left(\widehat{\Phi}_{\Lambda^\vee}(Y, \gamma) \right) + \sum_{\alpha,\beta,\tau=1}^n h^{\alpha\beta} Y_\alpha Y_\beta Y_\tau \frac{\partial}{\partial Y_\tau} \left(\widehat{\Phi}_{\Lambda^\vee}(Y, \gamma) \right).
\end{aligned} \tag{57}$$

Similarly, in terms of backward expression, we have

$$\begin{aligned}
& \sum_{\alpha,\beta=1}^n \frac{\partial}{\partial Y_\alpha} \left(Y_\alpha \frac{\partial}{\partial Y_\beta} \left(Y_\beta \widehat{\Phi}_{\Lambda^\vee}(Y, \gamma) \right) \right) = \sum_{\alpha,\beta=1}^n \frac{\partial^2}{\partial Y_\alpha \partial Y_\beta} \left(Y_\alpha Y_\beta \widehat{\Phi}_{\Lambda^\vee}(Y, \gamma) \right) \\
& - \sum_{\alpha=1}^n \frac{\partial}{\partial Y_\alpha} \left(Y_\alpha \widehat{\Phi}_{\Lambda^\vee}(Y, \gamma) \right), \\
& \sum_{\alpha,\beta,\tau=1}^n h^{\alpha\beta} Y_\alpha Y_\beta \frac{\partial}{\partial Y_\tau} \left(Y_\tau \widehat{\Phi}_{\Lambda^\vee}(Y, \gamma) \right) = \sum_{\alpha,\beta,\tau=1}^n h^{\alpha\beta} \frac{\partial}{\partial Y_\tau} \left(Y_\alpha Y_\beta Y_\tau \widehat{\Phi}_{\Lambda^\vee}(Y, \gamma) \right) \\
& - 2 \sum_{\alpha,\beta=1}^n h^{\alpha\beta} \left(Y_\alpha Y_\beta \widehat{\Phi}_{\Lambda^\vee}(Y, \gamma) \right) + \sum_{\alpha=1}^n h^{\alpha\alpha} (Y_\alpha^2 - Y_\alpha) \widehat{\Phi}_{\Lambda^\vee}(Y, \gamma).
\end{aligned} \tag{58}$$

By applying this lemma, from (59), we finally arrive at the following

Theorem 9. Assume that $\frac{1}{2} + \gamma$ is a zero of non-abelian zeta function $\widehat{\zeta}_{\mathbb{Q},n}(s)$ on the central line. Then, we have the following equivalent form of the average of Fokker-Planck equations over $\mathcal{M}_{\mathbb{Q},n}[1]$.

$$\begin{aligned}
(1) \quad & \int_{\mathcal{M}_{\mathbb{Q},n}[1]} \left(\sum_{\alpha,\beta=1}^n \frac{\partial}{\partial Y_\alpha} Y_\alpha \frac{\partial}{\partial Y_\beta} Y_\beta - n \sum_{\alpha=1}^n \frac{\partial}{\partial Y_\alpha} Y_\alpha \right. \\
& \quad \left. + 2\pi \left(\sum_{\alpha,\beta,\tau=1}^n h^{\alpha\beta} Y_\alpha Y_\beta \frac{\partial}{\partial Y_\tau} Y_\tau + \sum_{\alpha,\beta,\tau=1}^n h^{\alpha\beta} \frac{\partial}{\partial Y_\tau} Y_\alpha Y_\beta Y_\tau \right) \right. \\
& \quad \left. + 4\pi^2 \left(\sum_{\alpha,\beta=1}^n h^{\alpha\beta} Y_\alpha Y_\beta \right)^2 - 2\pi n \sum_{\alpha,\beta=1}^n h^{\alpha\beta} Y_\alpha Y_\beta + n^2 \left(\frac{1}{4} + \gamma^2 \right) \right) \widehat{\Phi}_{\Lambda^\vee}(Y, \gamma) \\
& \quad \quad \quad - \widehat{\kappa}_{\Lambda^\vee}(Y) \Big) d\mu = 0
\end{aligned} \tag{59}$$

$$\begin{aligned}
(2) \quad & \int_{\mathcal{M}_{\mathbb{Q},n}[1]} \left(\left(\sum_{\alpha,\beta=1}^n Y_\alpha Y_\beta \frac{\partial^2}{\partial Y_\alpha \partial Y_\beta} + \left(4\pi \sum_{\alpha,\beta=1}^n h^{\alpha\beta} Y_\alpha Y_\beta + (n+1) \right) \sum_{\alpha=1}^n Y_\alpha \frac{\partial}{\partial Y_\alpha} \right. \right. \\
& \quad \left. \left. + \left(2\pi \sum_{\alpha,\beta=1}^n h^{\alpha\beta} Y_\alpha Y_\beta + 1 \right)^2 - \left(2\pi \sum_{\alpha=1}^n h^{\alpha\alpha} Y_\alpha^2 + 1 \right) + n^2 \left(\frac{1}{4} + \gamma^2 \right) \right) \widehat{\Phi}_{\Lambda^\vee}(Y, \gamma) \right. \\
& \quad \quad \quad \left. - \widehat{\kappa}_{\Lambda^\vee}(Y) \right) d\mu = 0
\end{aligned} \tag{60}$$

$$\begin{aligned}
(3) \quad & \int_{\mathcal{M}_{\mathbb{Q},n}[1]} \left(\left(\left(\sum_{\alpha,\beta=1}^n \frac{\partial^2}{\partial Y_\alpha \partial Y_\beta} Y_\alpha Y_\beta - (n+1) \sum_{\alpha=1}^n \frac{\partial}{\partial Y_\alpha} Y_\alpha \left(1 + 4\pi \sum_{\beta,\tau=1}^n h^{\beta\tau} Y_\beta Y_\tau \right) \right. \right. \right. \\
& \quad \left. \left. + 2\pi \left(2\pi \left(\sum_{\alpha,\beta=1}^n h^{\alpha\beta} Y_\alpha Y_\beta \right)^2 - (n+2) \sum_{\alpha,\beta=1}^n h^{\alpha\beta} Y_\alpha Y_\beta + \sum_{\alpha=1}^n h^{\alpha\alpha} (Y_\alpha^2 - Y_\alpha) \right) \right) \widehat{\Phi}_{\Lambda^\vee}(Y, \gamma) \right. \\
& \quad \quad \quad \left. \left. + n^2 \left(\frac{1}{4} + \gamma^2 \right) \widehat{\Phi}_{\Lambda^\vee}(Y, \gamma) - \widehat{\kappa}_{\Lambda^\vee}(Y) \right) \right) d\mu = 0
\end{aligned} \tag{61}$$

The equations (59), (60), (61) are the variations of the equation

$$\begin{aligned}
& \int_{\mathcal{M}_{\mathbb{Q},n}[1]} \left(\left(\sum_{\alpha=1}^n \frac{\partial}{\partial Y_\alpha} Y_\alpha \right)^2 \left(B_2(Y) \cdot \widehat{\Phi}_{\Lambda^\vee}(Y, \gamma) \right) + \left(\sum_{\alpha=1}^n \frac{\partial}{\partial Y_\alpha} Y_\alpha \right) \left(B_1(Y) \widehat{\Phi}_{\Lambda^\vee}(Y, \gamma) \right) \right. \\
& \quad \quad \quad \left. + B_0(Y) \widehat{\Phi}_{\Lambda^\vee}(Y, \gamma) + \gamma^2 \widehat{\Phi}_{\Lambda^\vee}(Y, \gamma) - \widehat{\kappa}(Y) \right) d\mu = 0
\end{aligned} \tag{62}$$

for suitable functions $B_0(Y)$, $B_1(Y)$ and $B_2(Y)$ on the fibers of the fibration $\pi : \mathcal{T}_{\mathbb{Q},n} \rightarrow \mathcal{D}_1$ defined by the follows: for each $[\Lambda] \in \mathcal{M}_{\mathbb{Q},n}[1]$, the fiber of π over $[\Lambda]$ is the torus $\mathbb{T}_\Lambda = \mathbb{R}^n/\Lambda$. This may be viewed as an average on $\mathcal{M}_{\mathbb{Q},n}[1]$ form of the (backward and/or forward) *Fokker-Planck equation*, in dimension n .

5.2 Essential Projectively Flat Connection and Pro-Flat Sections

Consider the functions $\{\widehat{\Phi}_{\Lambda^\vee}(Y, \gamma)\}_\gamma$ parametrized by by the zeros of the rank n -non-abelian zeta function $\widehat{\zeta}_{\mathbb{Q},n}(s)$ on the central line. Smoothly depending on Y , we may

view $\{\widehat{\Phi_{\Lambda^\vee}(Y, \gamma)}\}_\gamma$ as an infinite family of smooth sections of the infinite dimensional Hilbert vector bundle $\mathcal{H}_{\widehat{\zeta}_n}$ on the universal family $\mathcal{T} \rightarrow \mathcal{M}_{\mathbb{Q}, n}[1]$. Apparently quite complicated, we believe that the Fokker-Planck equation reveals some natural geometric differential structure for this infinite dimensional Hilbert vector bundle $\mathcal{H}_{\widehat{\zeta}_n}$. To be more explicit, there should exist a Hilbert-metric-compatible projectively flat connection (modulo the integration over the base moduli space $\mathcal{M}_{\mathbb{Q}, n}[1]$) such that $\{\widehat{\Phi_{\Lambda^\vee}(Y, \gamma)}\}_\gamma$ are a base of the projectively flat connections again modulo the integration over the base moduli space $\mathcal{M}_{\mathbb{Q}, n}[1]$. If exists, we call such a connection an essential projectively flat connection and the above family a family of essential pro-flat section. This is the source of positivity which guarantees that all zeta zeros are on the central line.

6 Examples

6.1 Rank One

In the case, when $n = 1$, the above work is nothing but Armitage [1]. In fact, it is this work that we are modeled during our studies.

Since there is only one rank one lattice in \mathbb{R} given by $\mathbb{Z} \hookrightarrow \mathbb{R}$. Hence $\mathcal{M}_{\mathbb{Q}, 1}[1]$ consists of one point. This means that there is no family but only a single Fokker-Planck equation for the torus \mathbb{R}/\mathbb{Z} in classical sense.

In particular,

$$\widehat{\xi}_{\mathbb{Q}, 1}(s) = \widehat{\zeta}(s)$$

where $\widehat{\zeta}(s) := \pi^{-s/2} \Gamma(s/2) \zeta(s)$ is the complete Riemann zeta function.

Certainly, then the Riemann Hypothesis remains open. So, unlike the situation when $n \geq 2$, from the beginning, we cannot assume that the zeros of $\widehat{\zeta}_{\mathbb{Q}, 1}(s)$ has the form $\frac{1}{2} + i\gamma \in \frac{1}{2} + i\mathbb{R}$. The idea of Armitage [1] and Berry-Keating [?] is to use the Fokker-Planck equations in statistical mechanics, the Hamiltonians in quantum mechanics, and quantum harmonic oscillators to obtain positive definite operators whose eigen values coincide with the zeta zeros, so as to materialize an old idea of Bolyai-Hilbert.

To recover Armitage [1], we start with the heat equation for the theta series. Since $\mathcal{M}_{\mathbb{Q}, 1}[1]$ consists of one lattice $\mathbb{Z} \subset \mathbb{R}$, for which the dual lattice is itself, the metric matrix $H = (1)$ and the Laplacian operator is given by $\Delta_{\mathbb{Z}} = \frac{\partial^2}{\partial x^2}$, the heat equation (18) is specialized into

$$\begin{cases} \frac{1}{4\pi} \frac{\partial^2}{\partial x^2} \theta_{\mathbb{Z}}(x, t) = \frac{\partial}{\partial t} \theta_{\mathbb{Z}}(x, t) \\ \theta_{\mathbb{Z}}(x, 0) = \sum_{n \in \mathbb{Z}} e^{2\pi i n x} \end{cases} \quad (63)$$

where the theta function is given by $\theta_{\mathbb{Z}}(x, t) = \sum_{n \in \mathbb{Z}} e^{-\pi n^2 t + 2\pi i n x}$

Hence if we set $t = \exp(T)$ and $X = \exp(T/2) X$,

$$\Delta_{z, X} = \frac{\partial^2}{\partial X^2}, \quad \langle X, \frac{\partial}{\partial X} \rangle = X \frac{\partial}{\partial X} \quad \text{and} \quad \Theta_{\mathbb{Z}}(X, T) = \sum_{n \in \mathbb{Z}} e^{-\pi e^T n^2 + 2\pi i n e^{T/2} X}. \quad (64)$$

Hence, the corresponding heat equation (27) is specialized as the differential equation on the dual torus $\mathbb{T}_{\mathbb{Z}} \simeq \mathbb{R}^n / (e^{-T/2} \mathbb{Z})$ given by

$$\begin{cases} \left(\frac{\partial}{\partial T} - \Omega_{z, X} \right) (\Theta_{\mathbb{Z}}(X, T)) = 0, \\ \Theta_{\mathbb{Z}}(X, 0) = \theta_{\Lambda}(x, 1). \end{cases} \quad (65)$$

Here, $\Omega_{\Lambda^\vee, X}$ denotes the second order differential operator defined by

$$\Omega_{\Lambda^\vee, X} := \frac{1}{4\pi} \frac{\partial^2}{\partial X^2} + \frac{1}{2} X \frac{\partial}{\partial X}. \quad (66)$$

Consequently, (34) takes the form

$$2 \left(\widehat{\zeta}(s) - \frac{1}{s(s-1)} \right) = \left(\left(R \left(-\frac{s}{2}, \Omega_{\mathbb{Z}, X} \right) + R \left(-\frac{1-s}{2}, \Omega_{\mathbb{Z}, X} \right) \right) (\Theta_{\mathbb{Z}}(X, 0) - 1) \right) \Big|_{X=0}. \quad (67)$$

Consequently, if we assume that $\frac{1}{2} + i\gamma$ is a Riemann zero of $\widehat{\zeta}(s)$ on the central line, and set

$$\begin{aligned} \Phi_{\mathbb{Z}}^{(1)}(X, \gamma) &:= \frac{1}{2} \int_0^\infty \left(e^{(\frac{1}{4} + \frac{1}{2}i\gamma)T} + e^{(\frac{1}{4} - \frac{1}{2}i\gamma)T} \right) (\Theta_{\mathbb{Z}}(X, T) - 1) dT, \\ \kappa_{\mathbb{Z}}^{(1)}(X) &:= 2 \left(2\Omega_{\mathbb{Z}, X} + \frac{1}{2} \right) (\Theta_{\mathbb{Z}}(X, 0) - 1). \end{aligned} \quad (68)$$

and

$$\begin{aligned} \Phi_{\mathbb{Z}}(X, \gamma) &:= \Phi_{\mathbb{Z}}^{(1)}(X, \gamma) + \frac{1}{1/4 + \gamma^2} \exp(-\pi|X|^2), \\ \kappa_{\mathbb{Z}}(X) &:= \kappa_{\mathbb{Z}}^{(1)}(X) + \frac{1}{1/4 + \gamma^2} \left(\left(2\Omega_{\mathbb{Z}, X} + \frac{1}{2} \right)^2 + \gamma^2 \right) \exp(-\pi|X|^2), \end{aligned} \quad (69)$$

Hence we may introduce the special value of fat zeta function by

$$\begin{aligned} &\left(\left(2\Omega_{\mathbb{Z}, X} + \frac{1}{2} \right)^2 + \gamma^2 \right) \left(\widehat{\zeta} \left(X, \frac{1}{2} + i\gamma \right) - \frac{1}{1/4 + \gamma^2} (1 - \exp(-\pi|X|^2)) \right) \\ &= \left(\left(2\Omega_{\mathbb{Z}, X} + \frac{1}{2} \right)^2 + \gamma^2 \right) \Phi_{\mathbb{Z}}(X, \gamma) \\ &= \kappa_{\mathbb{Z}}(X). \end{aligned} \quad (70)$$

then Theorem 6 becomes

Theorem 10 (Armitage[[1]). *Assume that $\frac{1}{2} + i\gamma \in \frac{1}{2} + i\mathbb{R}$ is a zero of $\widehat{\zeta}(s)$. Then the function $\Phi_{\mathbb{Z}}(X, \gamma)$ is a solution of the following ‘average’ differential equation*

$$\left(\left(2\Omega_{\mathbb{Z}, X} + \frac{1}{2} \right)^2 + (\gamma)^2 \right) \Phi_{\mathbb{Z}}(X, \gamma) - \kappa_{\mathbb{Z}}(X) = 0 \quad (71)$$

and satisfies the following initial conditions

- (1) $\Phi_{\mathbb{Z}}(0, \gamma) = 0$,
- (2) $\frac{\partial}{\partial X} \Phi_{\mathbb{Z}}(0, \gamma) = \frac{\partial^3}{\partial X^3} \Phi_{\mathbb{Z}}(0, \gamma) = 0$;
- (3) $\lim_{|X| \rightarrow \infty} \left| \frac{\partial^j}{\partial X^j} \Phi_{\mathbb{Z}}(X, \gamma) \right| = O \left(\frac{1}{|X|} \right)$ for $0 \leq j \leq 4$.

Consequently, in terms of perturbations, we have the follows.

Suppose there would be a family of Hilbert spaces $\mathcal{H}_{\mathbb{Z}, 0}$ of functions defined on the torus $\mathbb{T}_{e^{-T/2\mathbb{Z}}}$ with an associated inner product (\cdot, \cdot) , containing both $\Phi_{\mathbb{Z}}(X, \gamma)$ and $\kappa_{\mathbb{Z}}(X)$ for all γ . Denote the real and the imaginary part of $\Phi_{\mathbb{Z}}(X, \gamma)$ by $\alpha_{\mathbb{Z}, \gamma}(X)$ and $\beta_{\mathbb{Z}, \gamma}(X)$. Then

$$\Phi_{\mathbb{Z}}(X, \gamma) = \alpha_{\mathbb{Z}, \gamma}(X) + i\beta_{\mathbb{Z}, \gamma}(X) \quad \text{and} \quad \Phi_{\mathbb{Z}}(X, \bar{\gamma}) = \alpha_{\mathbb{Z}, \gamma}(X) - i\beta_{\mathbb{Z}, \gamma}(X). \quad (72)$$

In the spaces $\mathcal{H}_{z,0}$, chose a family of function $\varphi_z(X)$ such that

$$\langle \beta_{z,\gamma}(X), \varphi_z(X) \rangle = 0 \quad \text{and} \quad \langle \alpha_{z,\gamma}(X), \varphi_z(X) \rangle \neq 0. \quad (73)$$

Hence if we define a family of projection operators

$$\begin{aligned} P_{z,\gamma} : \mathcal{H}_{z,0} &\rightarrow \langle \kappa_z(X) \rangle, \\ f &\mapsto -\frac{\langle f, \varphi_z(X) \rangle}{\langle \alpha_{z,\gamma}(X), \varphi_z(X) \rangle} \cdot \kappa_z(X) \end{aligned} \quad (74)$$

then

$$P_{z,\bar{\gamma}} = P_{z,\gamma}. \quad (75)$$

and

$$\left(\left(2\Omega_{\Lambda^\vee, X} + \frac{n}{2} \right)^2 + P_{\Lambda,\gamma} + (n\gamma)^2 \right) \Phi_{\Lambda^\vee}(X, \gamma) = 0. \quad (76)$$

if $\frac{1}{2} + \gamma$ is a zero of $\widehat{\zeta}(s)$ on the central line.

All in all, we have the following results on the Riemann zeros and the Fokker-Planck equations.

Theorem 11 (Armitage[1]). *Assume that $\frac{1}{2} + \gamma$ is a zero of non-abelian zeta function $\widehat{\zeta}(s)$ on the central line. Then, we have the following equivalent form of the Fokker-Planck equations.*

$$\begin{aligned} (1) \quad & \left(\frac{\partial}{\partial Y^2} - \frac{\partial}{\partial Y} Y + 2\pi \left(Y^2 \frac{\partial}{\partial Y} Y + \frac{\partial}{\partial Y} Y^3 \right) + 4\pi^2 Y^4 - 2\pi Y^2 + \left(\frac{1}{4} + \gamma^2 \right) \right) \widehat{\Phi}_z(Y, \gamma) \\ & - \widehat{\kappa_z(Y)} = 0 \end{aligned} \quad (77)$$

$$\begin{aligned} (2) \quad & \left(\left(Y^2 \frac{\partial^2}{\partial Y^2} + (4\pi Y^2 + 2) Y \frac{\partial}{\partial Y} + (Y^2 + 1)^2 - (2\pi Y^2 + 1) + \left(\frac{1}{4} + \gamma^2 \right) \right) \widehat{\Phi}_z(Y, \gamma) \right. \\ & \left. - \widehat{\kappa_z(Y)} = 0 \right) \end{aligned} \quad (78)$$

$$\begin{aligned} (3) \quad & \left(\left(\frac{\partial^2}{\partial Y^2} Y^2 - 2 \frac{\partial}{\partial Y} Y \left((1 + 4\pi) Y^2 \right) \right) \right. \\ & \left. + 2\pi \left(2\pi Y^4 - 3Y^2 + (Y^2 - Y) \right) + \left(\frac{1}{4} + \gamma^2 \right) \right) \widehat{\Phi}_z(Y, \gamma) - \widehat{\kappa_z(Y)} = 0 \end{aligned} \quad (79)$$

6.2 Rank Two

Rank two case is more complicated. It is the first case that the integration on the base moduli space is needed to obtain a uniformly stable theory.

Recall that, from the classical reduction, rank two lattices are parametrized by the fundamental domain \mathcal{D} of $\text{SL}_2(\mathbb{Z})$ in the upper half complex plane \mathcal{H} . More precisely, we have

$$\mathcal{D} := \left\{ z = x + iy \in \mathbb{C} : \begin{array}{l} x^2 + y^2 \geq 1, \quad -\frac{1}{2} \leq x < \frac{1}{2} \\ \& \text{if } x^2 + y^2 = 1, \quad x \leq 0 \end{array} \right\}$$

This can be used to give an explicit description of the moduli space $\mathcal{M}_{\mathbb{Q},2}[1]$ of semi-stable lattice of rank two and volume one. Indeed, by [5], there is a natural identification

$$\begin{aligned} \mathcal{M}_{\mathbb{Q},2}[1] &\simeq \mathcal{D}_1 := \{ z = x + iy \in \mathcal{D} : y \leq 1 \} \\ \Lambda = \left(\mathbb{Z}^2, \frac{1}{y} \begin{pmatrix} x^2 + y^2 & x \\ x & 1 \end{pmatrix} \right) &\mapsto \tau_\Lambda = x + iy \end{aligned}$$

In the sequel, we will simply use \mathcal{D}_1 for $\mathcal{M}_{\mathbb{Q},2}[1]$.

From the definition, the theta function $\theta_\Lambda(x, t)$ of Λ is given by

$$\theta_\Lambda(\mathbf{x}; t) = \theta_\Lambda(x_1, x_2; t) = \sum_{(m,n) \in \mathbb{Z}^2} e^{-\pi t \frac{(x^2+y^2)m^2 + 2xmn + n^2}{y^2}} e^{2\pi i(m x_1 + n x_2)}. \quad (80)$$

This is nothing but the theta function in terms of $\tau = x + yi$

$$\theta_\tau(\mathbf{x}; t) = \theta_\tau(x_1, x_2; t) = \sum_{(m,n) \in \mathbb{Z}^2} e^{-\pi t \frac{|m\tau + n|^2}{y^2}} e^{2\pi i(m x_1 + n x_2)}. \quad (81)$$

Similarly, we have

$$\theta_{\Lambda^\vee}(\mathbf{x}; t) = \sum_{(m,n) \in \mathbb{Z}^2} e^{-\pi t \frac{m^2 - 2xmn + (x^2+y^2)n^2}{y^2}} e^{2\pi i(m x_1 + n x_2)}. \quad (82)$$

which coincides with the theta function in terms of $\tau^\vee = -x + yi$

$$\theta_{\tau^\vee}(\mathbf{x}; t) = \sum_{(m,n) \in \mathbb{Z}^2} e^{-\pi t \frac{|m - n\tau|^2}{y^2}} e^{2\pi i(m x_1 + n x_2)}. \quad (83)$$

We notice that in fact

$$\theta_{\Lambda^\vee}(\mathbf{x}; t) = \theta_\Lambda(\mathbf{x}; t) \quad \text{and} \quad \theta_{\tau_\Lambda}(\mathbf{x}; t) = \theta_\tau(\mathbf{x}; t) \quad (84)$$

which can be obtained by simple changes of $\pm m$ and $\mp n$ within the pairs $(m, n) \in \mathbb{Z}^2$. This is not just an coincidence. Indeed, since the dual lattice Λ^\vee of Λ and its corresponding point in \mathcal{H} are

$$\Lambda^\vee = \left(\mathbb{Z}^2, \frac{1}{y} \begin{pmatrix} 1 & -x \\ -x & x^2 + y^2 \end{pmatrix} \right) \mapsto \tau_{\Lambda^\vee} = -\frac{x}{x^2 + y^2} + \frac{y}{x^2 + y^2}i =: \tilde{x} + i\tilde{y}$$

which is not a point of \mathcal{D} , the associated Laplacian operator on differential forms of $\mathbb{T}_{\Lambda^\vee}$ becomes

$$\Delta_{\Lambda^\vee, \mathbf{x}} = \frac{1}{y} \left(\frac{\partial^2}{\partial x_1^2} - 2x \frac{\partial^2}{\partial x_1 \partial x_2} + (x^2 + y^2) \frac{\partial^2}{\partial x_2^2} \right),$$

or better, in terms of τ_{Λ^\vee}

$$\begin{aligned} \Delta_{\tau^\vee} &= \frac{1}{\tilde{y}} \left(\frac{\partial^2}{\partial \tilde{x}_1^2} - 2\tilde{x} \frac{\partial^2}{\partial \tilde{x}_1 \partial \tilde{x}_2} + (\tilde{x}^2 + \tilde{y}^2) \frac{\partial^2}{\partial \tilde{x}_2^2} \right) \\ &= \frac{x^2 + y^2}{y} \left(\frac{\partial^2}{\partial x_1^2} + 2 \frac{x}{x^2 + y^2} \frac{\partial^2}{\partial x_1 \partial x_2} + \left(\frac{(-x)^2}{(x^2 + y^2)^2} + \frac{y^2}{(x^2 + y^2)^2} \right) \frac{\partial^2}{\partial x_2^2} \right) \\ &= \frac{1}{y} \left((x^2 + y^2) \frac{\partial^2}{\partial x_1^2} + 2x \frac{\partial^2}{\partial x_1 \partial x_2} + \frac{\partial^2}{\partial x_2^2} \right) = \Delta_\tau \end{aligned} \quad (85)$$

In addition, note that $\tau_{\Lambda^\vee} = \frac{-x + iy}{x^2 + y^2} = \frac{-\bar{\tau}_\Lambda}{\tau_\Lambda \bar{\tau}_\Lambda} = -\frac{1}{\tau_\Lambda}$ which is the fractional transform

image of τ_Λ under the element $\begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix}$. Therefore, τ_Λ and τ_{Λ^\vee} are $\text{SL}_2(\mathbb{Z})$ -equivalent.

This implies that both the lattices Λ and Λ^\vee correspond to the same moduli point $\tau = \tau_\Lambda \in \mathcal{D}_1$. Consequently, $\Lambda \simeq \Lambda^\vee$, or equivalently, the matrices $H_\Lambda = \frac{1}{y} \begin{pmatrix} x^2 + y^2 & x \\ x & 1 \end{pmatrix}$

and $H_{\Lambda^\vee} = \frac{1}{y} \begin{pmatrix} 1 & -x \\ -x & x^2 + y^2 \end{pmatrix}$ are $\text{SO}(2)$ -equivalent. For this reason, we will only use Λ instead of both Λ and its dual Λ^\vee .

Obviously, the heat equation form the theta function $\theta_\Lambda(x_1, x_2; t)$ becomes

$$\begin{cases} \frac{1}{4\pi} \Delta_{\Lambda, x} (\theta_\Lambda(x, t)) = \frac{\partial}{\partial t} (\theta_\Lambda(x, t)) \\ \theta_\Lambda(x, 0) = \sum_{(m, n) \in \mathbb{Z}^2} e^{2\pi i(m x_1 + n x_2)} \end{cases} \quad (86)$$

on the torus \mathbb{T}_Λ . In terms of $\tau_\Lambda \in \mathcal{D}_1$, we have

$$\begin{cases} \frac{1}{4\pi} \Delta_{\tau, x} (\theta_\tau(x, t)) = \frac{\partial}{\partial t} (\theta_\tau(x, t)) \\ \theta_\tau(x, 0) = \sum_{(m, n) \in \mathbb{Z}^2} e^{2\pi i(m x_1 + n x_2)} \end{cases} \quad (87)$$

Here, we have used the fact that

$$\Delta_{\Lambda, x} = \Delta_{\tau, x} = \frac{1}{y} \left((x^2 + y^2) \frac{\partial^2}{\partial x_1^2} + 2x \frac{\partial^2}{\partial x_1 \partial x_2} + \frac{\partial^2}{\partial x_2^2} \right).$$

To obtain the Fokker-Planck equation, we make a change of variables $t =: \exp(T)$, $(x_1, x_2) =: \exp(T/2)(X_1, X_2) =: \exp(T/2)X$. Accordingly, we define

$$\begin{aligned} \Delta_{\Lambda, X} := \Delta_{\tau, X} &:= \frac{1}{y} \left((x^2 + y^2) \frac{\partial^2}{\partial X_1^2} + 2x \frac{\partial^2}{\partial X_1 \partial X_2} + \frac{\partial^2}{\partial X_2^2} \right), \\ \left\langle X, \frac{\partial}{\partial X} \right\rangle &:= X_1 \frac{\partial}{\partial X_1} + X_2 \frac{\partial}{\partial X_2}. \end{aligned} \quad (88)$$

and

$$\begin{aligned} \Theta_\Lambda(X, T) &:= \Theta_\tau(X, T) := \theta_\Lambda(X \exp(T/2), \exp T) \\ &= \sum_{(m, n) \in \mathbb{Z}^2} e^{-\pi e^T \frac{|m\tau + n|^2}{y^2}} e^{2\pi i(m X_1 + n X_2) e^{T/2}}. \end{aligned} \quad (89)$$

Then, by (27), on the dual torus $\mathbb{T}_\Lambda \simeq \mathbb{R}^2 / (e^{-T/2} \Lambda)$, we have the following heat equation for the Theta function

$$\begin{cases} \left(\frac{\partial}{\partial T} - \Omega_{\tau, X} \right) (\Theta_\tau(X, T)) = 0, \\ \Theta_\tau(X, 0) = \theta_\tau(x, 1). \end{cases} \quad (90)$$

Here, $\Omega_{\tau, X}$ is a second order differential operator defined by

$$\Omega_{\tau, X} := \Omega_{\Lambda, X} := \frac{1}{4\pi} \Delta_{\tau, X} + \frac{1}{2} \left\langle X, \frac{\partial}{\partial X} \right\rangle. \quad (91)$$

Consequently, if we denote by $R(\kappa, \Omega_{\tau, X})$ the resolvent associated to $(-\kappa - \Omega_{\tau, X})^{-1}$, then for the rank 2 non-abelian zeta function $\widehat{\zeta}_{\mathbb{Q}, 2}(s)$,

$$\begin{aligned} &\left(\widehat{\zeta}_{\mathbb{Q}, 2}(s) - \frac{1}{s(s-1)} \text{vol}(\mathcal{D}_1) \right) \\ &= \int_{\mathcal{D}_1} \left((R(-s, \Omega_{\tau, X}) + R(-(1-s), \Omega_{\tau, X})) (\Theta_\tau(X, 0) - 1) \right) \Big|_{X=0} \frac{dx \wedge dy}{y^2}. \end{aligned} \quad (92)$$

Now, let $\frac{1}{2} + i\gamma$ be a zero of $\widehat{\zeta}_{\mathbb{Q}, 2}(s)$. We introduce the auxiliary functions $\Phi_{\tau, X}^{(1)}(X, \gamma) = \Phi_\Lambda^{(1)}(X, \gamma)$ and $\kappa_\tau^{(1)}(X) = \kappa_\Lambda^{(1)}(X)$ by

$$\begin{aligned} \Phi_\tau^{(1)}(X, \gamma) &:= \int_0^\infty \left(e^{(\frac{1}{2} + i\gamma)T} + e^{(\frac{1}{2} - i\gamma)T} \right) (\Theta_\tau(X, T) - 1) dT, \\ \kappa_\tau^{(1)}(X) &:= 4(2\Omega_{\tau, X} + 1) (\Theta_\tau(X, 0) - 1). \end{aligned} \quad (93)$$

and the fat zeta function $\widehat{\zeta}_{\mathbb{Q},2}\left(X, \frac{1}{2} + i\gamma\right)$ by

$$\begin{aligned} & \left((2\Omega_{\tau,X})^2 + 4\gamma^2 \right) \left(\widehat{\zeta}_{\mathbb{Q},2}\left(X, \frac{1}{2} + i\gamma\right) - \frac{\text{vol}(\mathcal{D}_1)}{1/4 + \gamma^2} \left(1 - \exp(-\pi|X|^2)\right) \right) \\ & := \int_{\mathcal{D}_1} \left((2\Omega_{\tau,X} + 1)^2 + 4\gamma^2 \right) \Phi_{\tau}(X, \gamma) \frac{dx \wedge dy}{y^2} \\ & = \int_{\mathcal{D}_1} \kappa_{\tau}(X) \frac{dx \wedge dy}{y^2}. \end{aligned} \quad (94)$$

Hence, by Theorem 6, we have the following

Theorem 12. *Assume that $\frac{1}{2} + i\gamma$ is a zero of non-abelian zeta function $\widehat{\zeta}_{\mathbb{Q},2}(s)$ on the central line. Then $\Phi_{\tau}(X, \gamma)$ is a solution of the following average differential equation*

$$\int_{\mathcal{D}_1} \left((2\Omega_{\tau,X} + 1)^2 + 4\gamma^2 \right) \Phi_{\tau}(X, \gamma) - \kappa_{\tau}(X) \frac{dx \wedge dy}{y^2} = 0 \quad (95)$$

and satisfies the following initial properties

$$\begin{aligned} (1) \quad & \int_{\mathcal{D}_1} \Phi_{\tau}(0, \gamma) \frac{dx \wedge dy}{y^2} = 0, \\ (2) \quad & \int_{\mathcal{D}_1} \frac{\partial}{\partial X_i} \Phi_{\tau}(0, \gamma) \frac{dx \wedge dy}{y^2} = \int_{\mathcal{D}_1} \frac{\partial^3}{\partial X_i^3} \Phi_{\tau}(0, \gamma) \frac{dx \wedge dy}{y^2} = 0; \\ (3) \quad & \lim_{|X| \rightarrow \infty} \left| \int_{\mathcal{D}_1} \frac{\partial^j}{\partial X_i^j} \Phi_{\tau}(X, \gamma) \frac{dx \wedge dy}{y^2} \right| = O\left(\frac{1}{|X|}\right) \text{ for } 0 \leq j \leq 4. \end{aligned}$$

Consequently, using the projection operator $P_{\Lambda, \gamma}$, we have

$$\int_{\mathcal{D}_1} \left((2\Omega_{\tau,X} + 1)^2 + P_{\Lambda, \gamma} + 4\gamma^2 \right) \Phi_{\tau}(X, \gamma) \frac{dx \wedge dy}{y^2} = 0. \quad (96)$$

provided that $\frac{1}{2} + \gamma i$ be a zero of $\widehat{\zeta}_{\mathbb{Q},2}(s)$, which is known on the central line.

All these then lead to the following

Theorem 13. *Assume that $\frac{1}{2} + i\gamma$ is a zero of non-abelian zeta function $\widehat{\zeta}_{\mathbb{Q},2}(s)$ on the central line. Then, we have the following equivalent form of the average of Fokker-Planck equations over $\mathcal{M}_{\mathbb{Q},2}[1]$.*

(1) *The Fokker-Planck equation over the universal family $\pi : \mathcal{T} \rightarrow \mathcal{D}_1$ of tori is given by*

$$\begin{aligned} & \int_{\mathcal{D}_1} \left(\left(\frac{\partial}{\partial Y_1} Y_1 \frac{\partial}{\partial Y_1} Y_1 + 2 \frac{\partial}{\partial Y_1} Y_1 \frac{\partial}{\partial Y_2} Y_2 + \frac{\partial}{\partial Y_2} Y_2 \frac{\partial}{\partial Y_2} Y_2 - 2 \frac{\partial}{\partial Y_1} Y_1 - 2 \frac{\partial}{\partial Y_2} Y_2 \right. \right. \\ & \quad + \frac{2\pi}{y} \left((x^2 + y^2) \left(Y_1^2 \frac{\partial}{\partial Y_1} Y_1 + \frac{\partial}{\partial Y_1} Y_1^3 \right) + (x^2 + y^2) \left(Y_1^2 \frac{\partial}{\partial Y_2} Y_2 + \frac{\partial}{\partial Y_2} Y_1^2 Y_2 \right) \right. \\ & \quad \quad \left. \left. + 2x \left(Y_1 Y_2 \frac{\partial}{\partial Y_1} Y_1 + \frac{\partial}{\partial Y_1} Y_1^2 Y_2 \right) + \left(Y_1 Y_2 \frac{\partial}{\partial Y_2} Y_2 + \frac{\partial}{\partial Y_2} Y_1 Y_2^2 \right) \right) \right. \\ & \quad \left. + Y_2^2 \left(\frac{\partial}{\partial Y_1} Y_1 + \frac{\partial}{\partial Y_1} Y_1 \right) + \left(Y_2^2 \frac{\partial}{\partial Y_2} Y_2 + \frac{\partial}{\partial Y_2} Y_2^3 \right) \right) \\ & \quad + \frac{4\pi^2}{y^2} \left((x^2 + y^2) Y_1^2 + 2x Y_1 Y_2 + Y_2^2 \right)^2 - \frac{4\pi}{y} \left((x^2 + y^2) Y_1^2 + 2x Y_1 Y_2 + Y_2^2 \right) \\ & \quad \left. + 4 \left(\frac{1}{4} + \gamma^2 \right) \right) \widehat{\Phi}_{\tau}(Y, \gamma) - \widehat{\kappa}_{\tau}(Y) \frac{dx \wedge dy}{y^2} = 0 \end{aligned} \quad (97)$$

(2) The forward Fokker-Planck equation over the universal family $\pi : \mathcal{T} \rightarrow \mathcal{D}_1$ of tori is given by

$$\begin{aligned} & \int_{\mathcal{D}_1} \left(\left(Y_1^2 \frac{\partial^2}{\partial Y_1^2} + 2Y_1 Y_2 \frac{\partial^2}{\partial Y_1 \partial Y_2} + Y_2^2 \frac{\partial^2}{\partial Y_2^2} \right) \right. \\ & \quad + \left. \left(\frac{4\pi}{y} \left((x^2 + y^2) Y_1^2 + 2x Y_1 Y_2 + Y_2^2 \right) + 3 \right) \left(Y_1 \frac{\partial}{\partial Y_1} + Y_2 \frac{\partial}{\partial Y_2} \right) \right. \\ & \quad + \left. \left(\left(\frac{4\pi}{y} \left((x^2 + y^2) Y_1^2 + 2x Y_1 Y_2 + Y_2^2 \right) + 1 \right)^2 - \left(\frac{2\pi}{y} \left((x^2 + y^2) Y_1^2 + Y_2^2 \right) + 1 \right) \right. \right. \\ & \quad \quad \left. \left. + 4 \left(\frac{1}{4} + \gamma^2 \right) \right) \widehat{\Phi}_\tau(Y, \gamma) - \widehat{\kappa}_\tau(Y) \right) \frac{dx \wedge dy}{y^2} = 0 \end{aligned} \quad (98)$$

(3) The backward Fokker-Planck equation over the universal family $\pi : \mathcal{T} \rightarrow \mathcal{D}_1$ of tori is given by

$$\begin{aligned} & \int_{\mathcal{D}_1} \left(\left(\left(\frac{\partial^2}{\partial Y_1^2} Y_1^2 + 2 \frac{\partial^2}{\partial Y_1 \partial Y_2} Y_1 Y_2 + \frac{\partial^2}{\partial Y_2^2} Y_2^2 \right) \right. \right. \\ & \quad \left. \left. - 3 \left(\frac{\partial}{\partial Y_1} Y_1 + \frac{\partial}{\partial Y_2} Y_2 \right) \left(1 + \frac{4\pi}{y} \left((x^2 + y^2) Y_1^2 + 2x Y_1 Y_2 + Y_2^2 \right) \right) \right) \right. \\ & \quad + 2\pi \left(\frac{2\pi}{y^2} \left((x^2 + y^2) Y_1^2 + 2x Y_1 Y_2 + Y_2^2 \right)^2 - \frac{4}{y} \left((x^2 + y^2) Y_1^2 + 2x Y_1 Y_2 + Y_2^2 \right) \right. \\ & \quad \left. \left. + \left(Y_1^2 + Y_2^2 \right) - (Y_1 + Y_2) \right) \right) \\ & \quad \left. + 4 \left(\frac{1}{4} + \gamma^2 \right) \right) \widehat{\Phi}_\tau(Y, \gamma) - \widehat{\kappa}_\tau(Y) \right) \frac{dx \wedge dy}{y^2} = 0 \end{aligned} \quad (99)$$

Apparently very complicated, it is clear that these different equations are of the second order and elliptic on the fiber direction of the fibration $\pi : \mathcal{T} \rightarrow \mathcal{D}_1$, which becomes stable under an average on the base \mathcal{D}_1 . In this sense, $1 + 4\gamma^2$ appears as the eigenvalue for the eigenfunction $wh\Phi_\tau(Y, \gamma)$ of the corresponding second order elliptic operators.

7 Moduli Spaces of Abelian Varieties and Curves

There are natural generalizations of non-abelian zeta functions to the zeta functions of the special linear groups SL_n and their maximal parabolic subgroups P_{n_1, n_2, \dots, n_k} determined by ordered partitions $n = n_1 + n_2 + \dots + n_k$, and more generally, to the zeta functions of pairs consisting of a split reductive group and its maximal parabolic subgroup. One of the main purpose of [5] is to construct these zeta functions and study their basic properties. Best of all, we establish there a weak version for the Riemann hypothesis for all these zeta functions when working over the field of rationals, claiming that all but finitely many zeros of these zeta functions lie on the central line $\Re(s) = \frac{1}{2}$, provided that the rank of all its simple factors of G is at least one of course (hence the original Riemann hypothesis for the Riemann zeta function is not included). Therefore, the discussions above have their companions the lattices Λ replacing by the lattices in $\text{Lie}(G)_{\mathbb{R}}$ associated to what we call compatible arithmetical G -torsors and the space $\mathcal{M}_{\mathbb{Q}, n}$ by the moduli spaces of semi-stable arithmetic G -torsors of slope zero

In addition, probably more challenging, we can work over the moduli spaces $\mathcal{A}_{\mathbb{Q}, n}$ of abelian varieties whose structural lattices are semi-stable. This new moduli space is a compact subspaces of the moduli spaces of polarized abelian varieties of dimension n . In this acse, certainly, the reductive group G should be taken to be the symplectic

group Sp_{2n} . Similar structure as above should exist. What are the implications of these new structures

There are too many questions to be asked. For one, we may yet further construct a new type of zeta functions by working over the moduli spaces $\mathcal{C}_{\mathbb{Q},g}$ of regular curves of genus g for which the lattices of their Jacobians are assumed to be semi-stable. The moduli spaces $\mathcal{C}_{\mathbb{Q},g}$ may be viewed as a subspaces in the $\mathcal{A}_{\mathbb{Q},n}$, via Jacobean embeddings, or better, subspaces of the moduli spaces of curves \mathcal{M}_p . With the kdV equation characterizations for the theta functions of curves, we expect that there is a much more refined structures here as well.

A Fokker-Planck Equation

A.1 One Variable

The single variable *Fokker-Planck equation* with time-independent drift and diffusion coefficients $D^{(1)}(x)$ and $D^{(2)}(x)$ is given by

$$\frac{\partial}{\partial t}W(x,t) = L_{\mathrm{FP}}W(x,t) \quad (100)$$

where the Fokker-Planck (differential) operator is defined by

$$L_{\mathrm{FP}} = -\frac{\partial}{\partial x}D^{(1)}(x) + \frac{\partial^2}{\partial x^2}D^{(2)}(x). \quad (101)$$

In many cases, the Fokker-Planck equation (100) satisfied by the distribution function $W(x,t)$ characterizes the time-dependent distribution function.

Example 1. When a small particle of mass m of the velocity $v(t)$ is immersed in a fluid, because of the collisions between the molecules of the fluid and the particle, the friction force arises. Accordingly, the momentum of the particle is transferred to the molecules of the fluid and the velocity of the particle gradually decreases to zero. To describe this phenomenon, we first assume that the mass of the particle is large enough so that its velocity due to thermal fluctuations is negligible. Then the resulting friction force is, accordingly to Stokes's law, given by

$$F_c(t) = -\alpha v(t) \quad (102)$$

where α is a constant depending on the material. Since this friction force can be written in terms of acceleration $\dot{v}(t)$ by

$$F_c(t) = m\dot{v}(t). \quad (103)$$

All these then leads to the equation of motions

$$m\dot{v}(t) + \alpha v(t) = 0. \quad (104)$$

Therefore, the motion is completely determined by the relation

$$v(t) = v(0) e^{-\frac{\alpha}{m}t}. \quad (105)$$

However, when the mass becomes very small, the thermal velocity v_{th} is observed. Since, by the equipartition law, the mean energy of the particle is given by

$$\frac{1}{m}\langle v^2 \rangle = \frac{1}{2}\gamma T \quad (106)$$

where γ is the Boltzmann constant and T is the temperature.¹ Hence, the thermal velocity is calculated by

$$v_{\text{th}} = \sqrt{\langle v^2 \rangle} = \sqrt{\frac{\gamma T}{m}}. \quad (107)$$

In particular, the smaller the mass m is, the more the thermal velocity v_{th} becomes observable. As a result, for small mass particles, their velocities cannot be characterized by (104).

Nevertheless, no matter how small, when the mass of the particle is still bigger than the mass of the molecules, (104) can be modified to correct thermal energy by adding a fluctuating force $F_{\text{fl}}(t)$, the so-called *Langevin force*. That is to say, instead of (103), we get

$$-m\dot{v}(t) = F(t) = F_c(t) + F_{\text{fl}}(t) = -\alpha v(t) + F_{\text{fl}}(t) \quad (108)$$

In particular, if we denote the the Langevin force by $\Gamma(t) := \frac{1}{m} F_{\text{fl}}(t)$, the relation (104) is modified to

$$\dot{v}(t) + \frac{\alpha}{m} v(t) = \Gamma(t). \quad (109)$$

Since the force $F_{\text{fl}}(t)$ is a stochastic or a random force, their properties are only given in the average. In terms of the Langevin force $\Gamma(t)$, since the motion of the average velocity $\langle v(t) \rangle$ should be given by (103), we may assume that both the average of $\Gamma(t)$ and the average of the correlation of two Langevin forces for time differences $t' - t$ which are larger than the duration time τ_0 of a collision are zero, i.e.

$$\langle \Gamma(t) \rangle = 0 \quad \text{and} \quad \langle \Gamma(t)\Gamma(t') \rangle = 0 \quad \forall |t - t_0| \geq \tau_0. \quad (110)$$

In physics, the second relation then finally leads to the relation that

$$\langle \Gamma(t)\Gamma(t') \rangle = 2\frac{\alpha\gamma T}{m^2} \delta(t - t') \quad (111)$$

where δ denotes the Dirac δ -distribution.

Moreover, since $\Gamma(t)$, varying from system to system in the ensemble, is a stochastic quantity, and so is the velocity. Its distribution function, or the same, the probability density, $W(v, t)$ is known to satisfies the one-variable Fokker-Planck equation

$$\frac{\partial W}{\partial t} = \frac{\alpha}{m} \frac{\partial (vW)}{\partial v} + \frac{\alpha\gamma T}{m^2} \frac{\partial^2 W}{\partial v^2}. \quad (112)$$

Here, being a probability density, we have, for any function $g(v)$ of v ,

$$\langle g(v(t)) \rangle = \int_{-\infty}^{\infty} g(v)W(v, t) dv. \quad (113)$$

The Fokker-Planck equation (112) is one of the simplest forms, with constant coefficients $\frac{\alpha}{m}$ and $\frac{\alpha\gamma T}{m^2}$ of the first order and second order partial differentials, respectively. General Fokker-Planck equationa are taken the form

$$\frac{\partial W(x, t)}{\partial t} = \left(-\frac{\partial}{\partial x} D^{(1)}(x) + \frac{\partial^2}{\partial x^2} D^{(2)}(x) \right) W(x, t). \quad (114)$$

For our own conference, we call (114) a forward Fokker-Planck equation. Similarly, by a backward Fokker-Planck equation, we mean that of the form

$$\frac{\partial W(x, t)}{\partial t} = \left(-D^{(1)}(x) \frac{\partial}{\partial x} + D^{(2)}(x) \frac{\partial^2}{\partial x^2} \right) W(x, t). \quad (115)$$

¹For the meaning of $\langle \rangle$, please refer (??).

To see how these equations arise, let us consider a general Langevin equation on one stochastic variable ξ

$$\dot{\xi} = h(\xi, t) + g(\xi, t)\Gamma(t) \quad (116)$$

such that the Langevin force (with multiplicative noise) is bounded by the conditions that

$$\langle \Gamma(t) \rangle = 0 \quad \text{and} \quad \langle \Gamma(t)\Gamma(t') \rangle = 2\delta(t-t'). \quad (117)$$

Here multiplicative refers to the fact that $g(\xi, t)$ is not a constant in which case, we get an additive noise. But such a noise can be easily treated, since the difference between the multiplicative and additive noises are not that significant because, a simple change of variables would imply

$$\dot{\eta} = \frac{\dot{\xi}}{g} = \frac{h}{g} + \Gamma(t). \quad (118)$$

Assume (116). If we introduce the Kramer-Moyer expansion coefficients $D^{(n)}(x, t)$ by

$$D^{(n)}(x, t) := \frac{1}{n!} \lim_{\tau \rightarrow 0} \frac{\langle [\xi(t+\tau) - x]^n \rangle}{\tau} \Big|_{\xi(t)=x}, \quad (119)$$

then

$$D^{(n)}(x, t) = \begin{cases} h(x, t) + \frac{\partial g(x, t)}{\partial x} g(x, t) & n = 1, \\ g^2(x, t) & n = 2 \\ 0 & n \geq 3 \end{cases} \quad (120)$$

In physics term, $h(x, t)$ is called the deterministic drift and the so-called noise-induced drift contained in $D^{(1)}$ is defined by

$$D_{\text{ind}}^{(1)} := \frac{\partial g(x, t)}{\partial x} g(x, t) = \frac{1}{2} \frac{\partial}{\partial x} D^{(2)}(x, t). \quad (121)$$

A.2 Several Variables

Let $\xi = (\xi_1, \dots, \xi_n)$ be N stochastic variables satisfying the following system of Langevin equations

$$\begin{cases} \dot{\xi}_1 = h_1(\xi) + \sum_{\beta=1}^n g_{1\beta}(\xi, t)\Gamma_\beta(t) \\ \dots\dots\dots\dots\dots\dots\dots\dots\dots\dots\dots\dots\dots\dots\dots\dots\dots\dots \\ \dot{\xi}_n = h_n(\xi) + \sum_{\beta=1}^n g_{n\beta}(\xi, t)\Gamma_\beta(t) \end{cases} \quad (122)$$

subject the following constrains

$$\langle \Gamma_i(t) \rangle = 0 \quad \text{and} \quad \langle \Gamma_i(t)\Gamma_j(t') \rangle = 2\delta_{ij} \delta(t-t') \quad (123)$$

As in (119), we introduce the Kramer-Moyal coefficients

$$D_{i_1 \dots i_\nu}(x, t) := D_{i_1 \dots i_\nu}^{(\nu)}(x, t) := \frac{1}{\nu!} \lim_{\tau \rightarrow 0} \frac{1}{\tau} \langle [\xi_{i_1}(t+\tau) - x_{i_1}] \dots [\xi_{i_\nu}(t+\tau) - x_{i_\nu}] \rangle. \quad (124)$$

In particular, the drift and diffusion coefficients are hence given by

$$\begin{cases} D_i(x, t) = D_i^{(1)}(x, t) := \lim_{\tau \rightarrow 0} \frac{\langle \xi_i(t+\tau) - x_i \rangle}{\tau} \Big|_{\xi_k(t)=x_k} & k = 1, 2, \dots, n \\ D_{ij}(x, t) = D_{ij}^{(2)}(x, t) := \frac{1}{2} \lim_{\tau \rightarrow 0} \frac{\langle [\xi_i(t+\tau) - x_i][\xi_j(t+\tau) - x_j] \rangle}{\tau} \Big|_{\xi_k(t)=x_k} & k = 1, 2, \dots, n \end{cases} \quad (125)$$

It is not too difficult then to conclude that

$$\begin{cases} D_i(x, t) &= h_i(x, t) + \sum_{k,\beta=1}^n g_{ki}(x, t) \frac{\partial}{\partial x_k} g_{ij}(x, t) \\ D_{ij}(x, t) &= \sum_{k=1}^n g_{ik}(x, t) g_{jk}(x, t) \\ D_{i_1 \dots i_\nu}(x, t) &= 0 \quad (\nu \geq 3) \end{cases} \quad (126)$$

Similar as in one variable case, in fact, the drift and diffusion coefficients determine the Langevin forces, deterministic drifts $h_i(x, t)$ and noise-induced drifts $g_{ij}(x, t)$. Indeed, if we set D be matrix $D = (D_{ij})$, then

$$\begin{cases} g_{ij} = (D^{\frac{1}{2}})_{ij}, \\ h_i = D_i - \sum_{k,\beta=1}^n (D^{\frac{1}{2}})_{kj} \frac{\partial}{\partial x_k} (D^{\frac{1}{2}})_{ij} \end{cases} \quad (127)$$

Here, we have use the fact that $D = (D_{ij})$ is a the symmetric positive definite matrix, and hence its square root makes sense (by the taking positive of the square of its eigenvalues).

Since the Langevin equation (116) with δ -correlated Langevin forces is a Markov process, namely, its conditional probability at time t_n depends only on the variable $\xi(t_{n-1}) = x_{n-1}$ at the next earlier time, i.e.

$$P(x_n, t_n | x_{n-1}, t_{n-1}; \dots; x_1, t_1) = P(x_n, t_n | x_{n-1}, t_{n-1}) W_{n-1}(x_{n-1}, t_{n-1}; \dots; x_1, t_1), \quad (128)$$

we get the following Fokker-Planck equation, or the same, the forward Kolmogorov equation

$$\left(\frac{\partial}{\partial t} - L_{\text{FP}}(x, t) \right) W(x, t) = 0 \quad (129)$$

where

$$L_{\text{FP}}(x, t) = - \sum_{\alpha=1}^N \frac{\partial}{\partial x_i} D_i(x, t) + \sum_{\alpha,\beta=1}^n \frac{\partial^2}{\partial x_i \partial x_j} D_{ij}(x, t). \quad (130)$$

In general, the Langevin equation (116) does not give rise to a Markov process, it is well known that by introducing new random variables, non-Markovian process can be deduced from a Markov one. In this sense, the Fokker-Planck equation is really a general law. In fact these equations play very important roles in various branches of disciplines.

The point we want to make in this paper is that while the above discussions work well for a fixed ensemble for which the systems lie on, when there is a family of ensembles, the Langevin forces does not act in a simply way as if the ensembles involved would be totally independent. Contrary to this, the family of Langevin forces interact with each other within the family so that only after taking an average on the base space over which the ensembles form a family. Indeed, as what we observe in the main text, for the Fokker-Planck equation to stand on, there should be an average on the base space in force. In other words, there is no a single Langevin equation which dominates each ensemble in a family, but a global type of relations on taking the averages of family Langevin forces over the parametrized spaces. We call such an equation a global average force equations. This is the essence of our current work.

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